



## An extremely high altitude plume seen at Mars morning terminator

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This presentation is dedicated to the memory of Donald C. Parker (1939 - 22 Feb.2015), a friend, an exceptional planetary observer, coauthor of this research and regular contributor to the GCP research activities.



# MARS HIGH PLUME → 'AN EXPLOSIVE COCKTAIL'

- Mars (First GCP paper on that subject)
- Amateur observations in the spacecraft era (3 rovers and 5 orbiters active)
- Published in NATURE (Editorial comment – promotion)

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### An extremely high-altitude plume seen at Mars' morning terminator

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*Nature* 518, 525–528 (26 February 2015) | doi:10.1038/nature14162

Last updated: 24 May 2015 4:59:12 EDT

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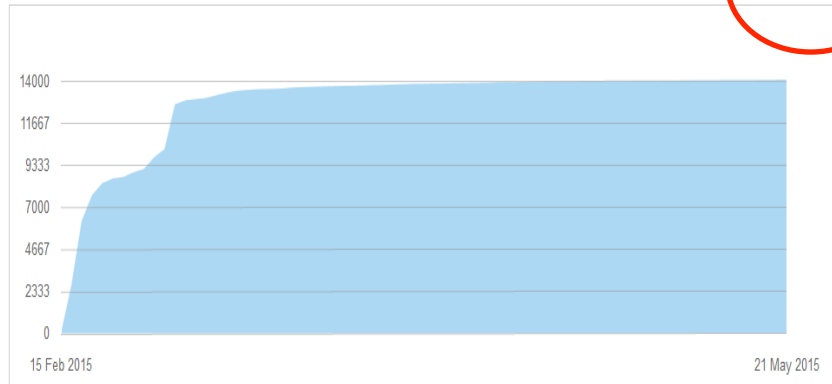
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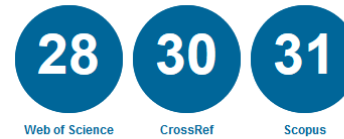
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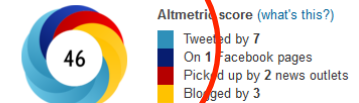
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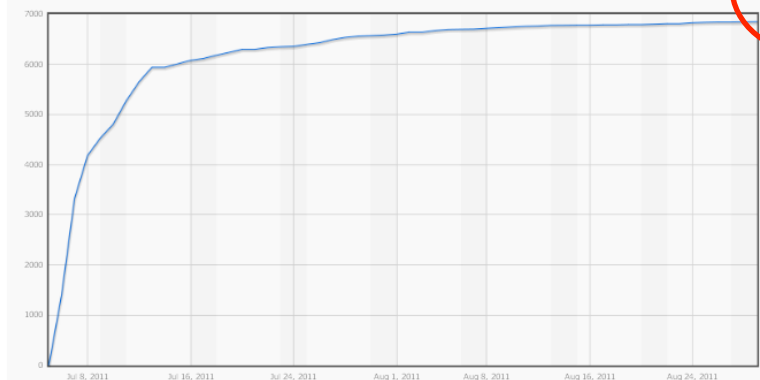
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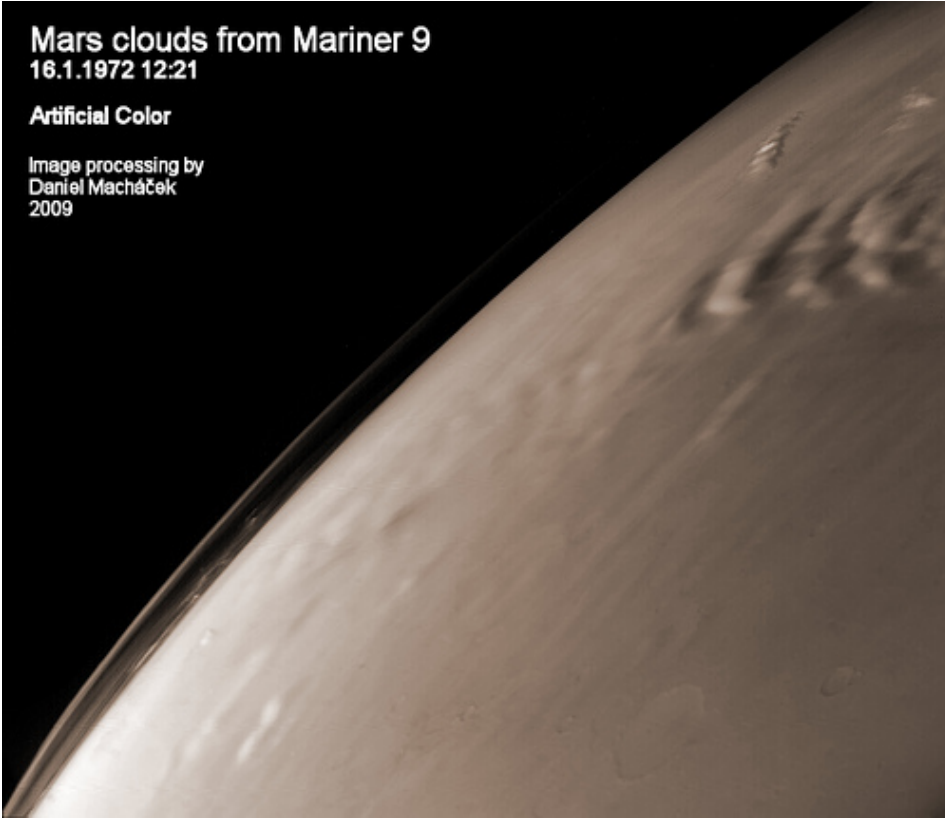


## Projected aerosols at Martian limb: Detached layers - I

Mars clouds from Mariner 9  
16.1.1972 12:21

Artificial Color

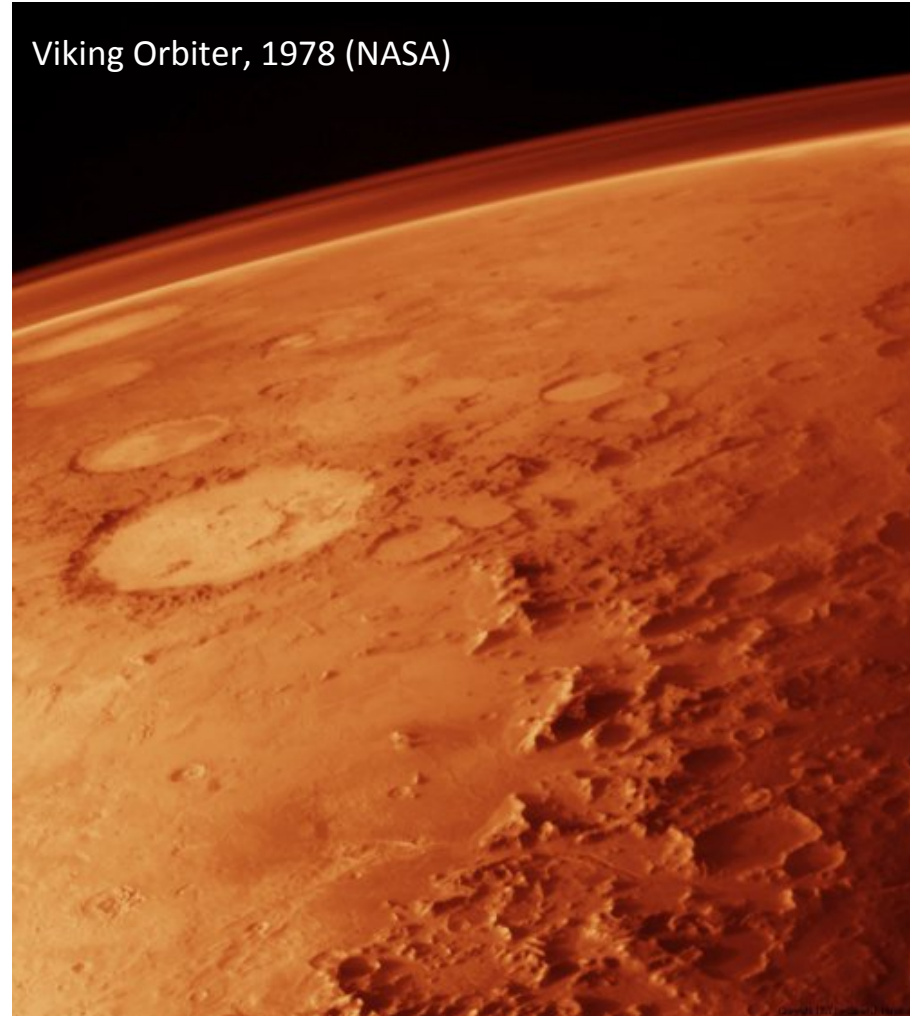
Image processing by  
Daniel Macháček  
2009



$Z(\text{max}) \sim 40 - 60 \text{ km}$

Anderson and Leovy, J. Atmos. Sci. (1978)

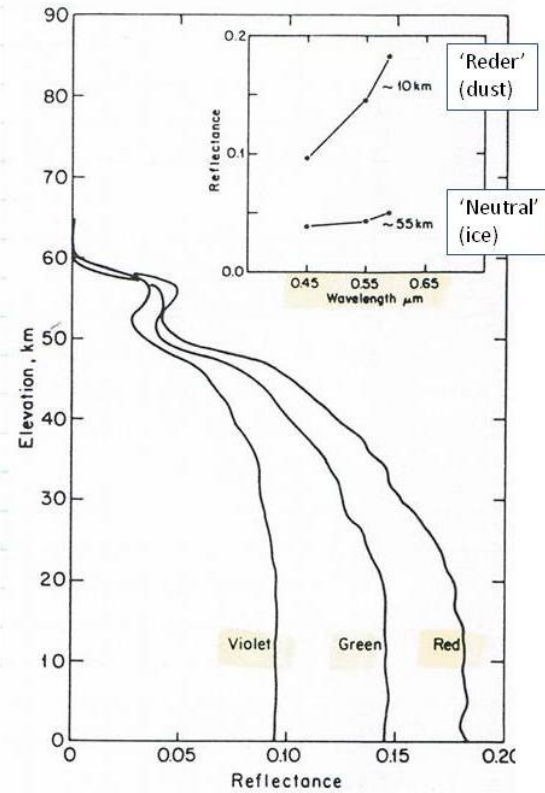
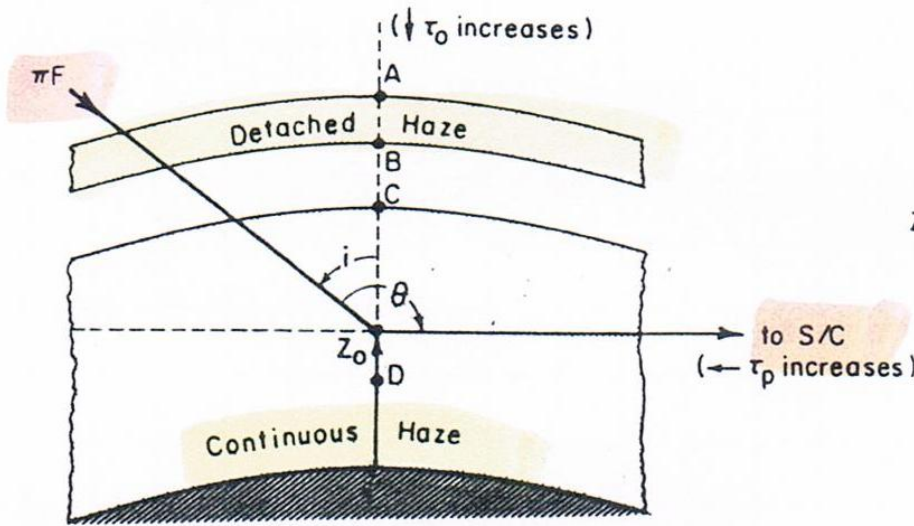
Viking Orbiter, 1978 (NASA)



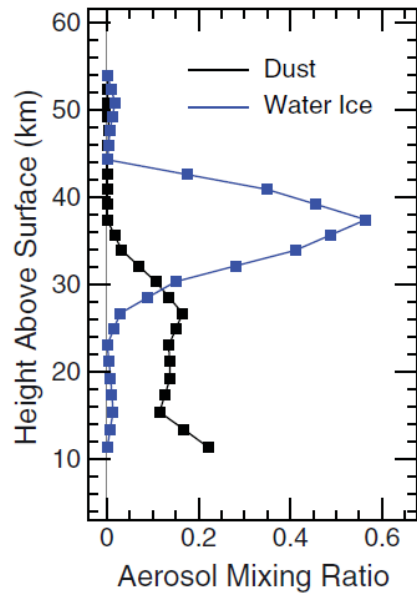
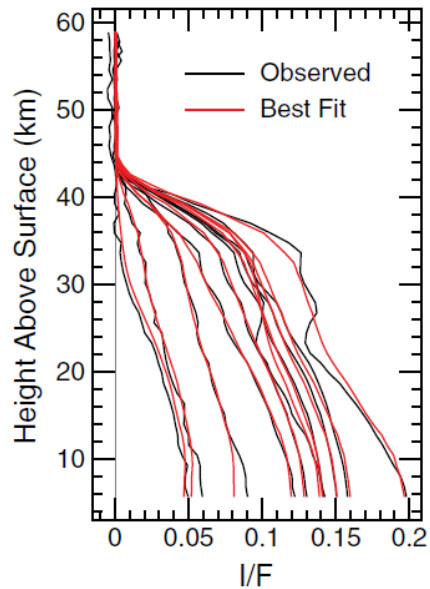
$Z(\text{max}) \sim 60 \text{ km}$

Jaquin et al., Icarus (1986)

# Projected aerosols at Martian limb: Detached layers - II



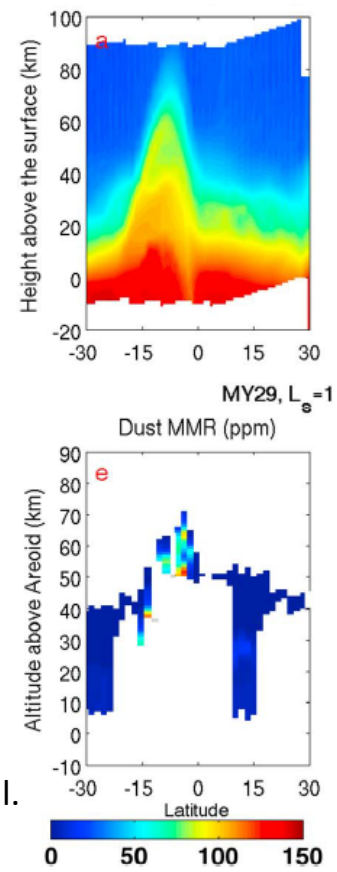
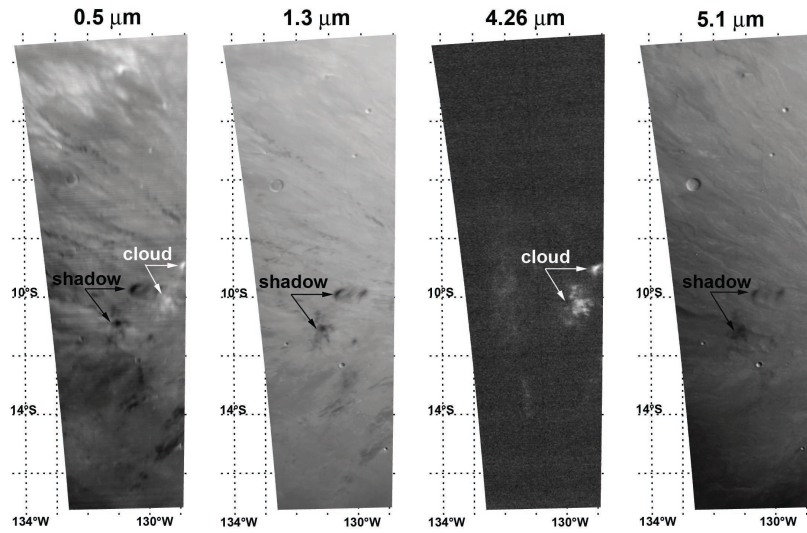
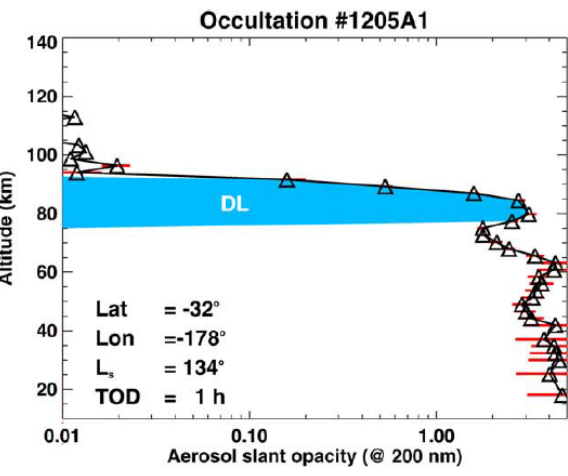
Jaquin et al., Icarus (1986)



Smith et al., JGR (2013)

Continuous haze = dust (mixed water-ice)  
 Detached haze = H<sub>2</sub>O-ice  
 Haze = Dust + H<sub>2</sub>O-ice  
 Z(max) Dust ~ 50 km  
 Z(max) H<sub>2</sub>O-ice ~ 60 km  
 Aerosols :  $r_{\text{eff}} = 0.5\text{-}3 \mu\text{m} \rightarrow r_{\text{eff}}(z)$   
 H<sub>2</sub>O-ice hexagonal cristals

# Clouds at High altitude



CO<sub>2</sub>-ice clouds  
Montmessin et al., Icarus (2006)

CO<sub>2</sub>-ice clouds up to 100 km  
Maattanen et al., Icarus (2010)

Dust  
Heavens et al.  
GRL (2015)

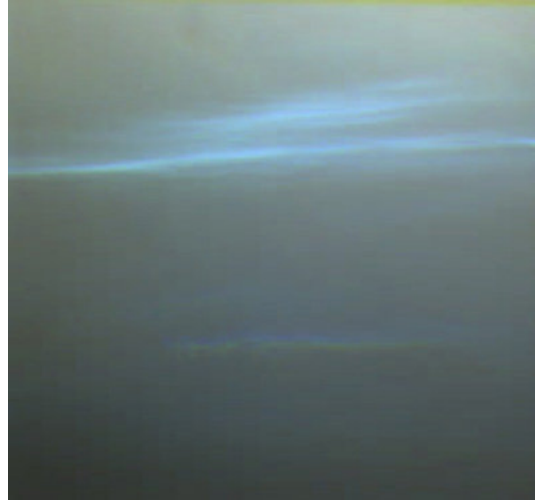
- Imaging & Photopolarimeter at limb
- Limb solar occultation (Phobos2)
- Star occultation (SPICAM/MEX)
- Limb spectra (CRISM/MRO)
- Radiance scans (TES/MGS; MCS/MRO)
- Imaging & shadows on disk (OMEGA-HRSC/MEX)
- Thermal (THEMIS/MOdysey)

- High altitude ice-clouds:**
- H<sub>2</sub>O-ice (Z (max) = 70 -90 km;  $r_{\text{eff}} = 0.2-0.3 \mu\text{m}$ )
- CO<sub>2</sub>-ice clouds (Z = 90-100 km;  $r_{\text{eff}} = 0.1 \mu\text{m}$ )
- Extreme Detached Dust Layers (EDDL)**
- Z(top) ~ 80 km (over volcanos)

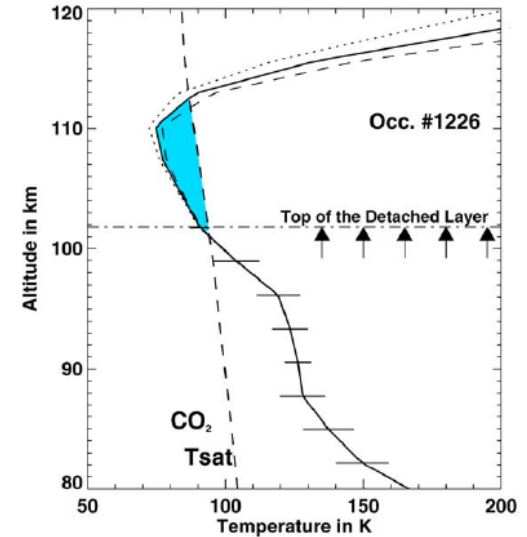
# Origin of high altitude aerosols



H<sub>2</sub>O-ice clouds. Phoenix.  
 Lat = 68.2°N  
 Z(land) = -4.1 km.  
 Whiteway et al., Science (2009)



CO<sub>2</sub>-ice clouds .Pathfinder.  
 Lat = 19.1°N (Z > 70 km).  
 Smith et al., Science (1997)

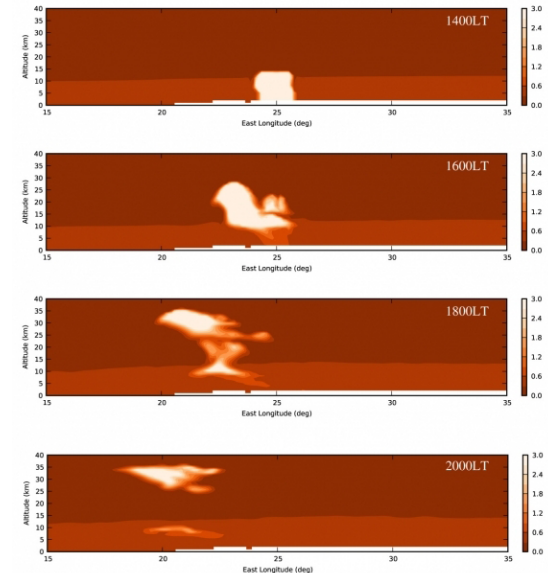


Montmessin et al., Icarus (2006)

**\* Mainly in the Equatorial-Tropical band (~ 30°N-30°S)**  
**Topography and season.** L (horizontal) ~ 100 – 1,000 km

**\* High altitude H<sub>2</sub>O-ice & CO<sub>2</sub>-ice clouds:**  
 Condensation on supercooled Cold air: Dynamics & transported by waves (tides, gravity)

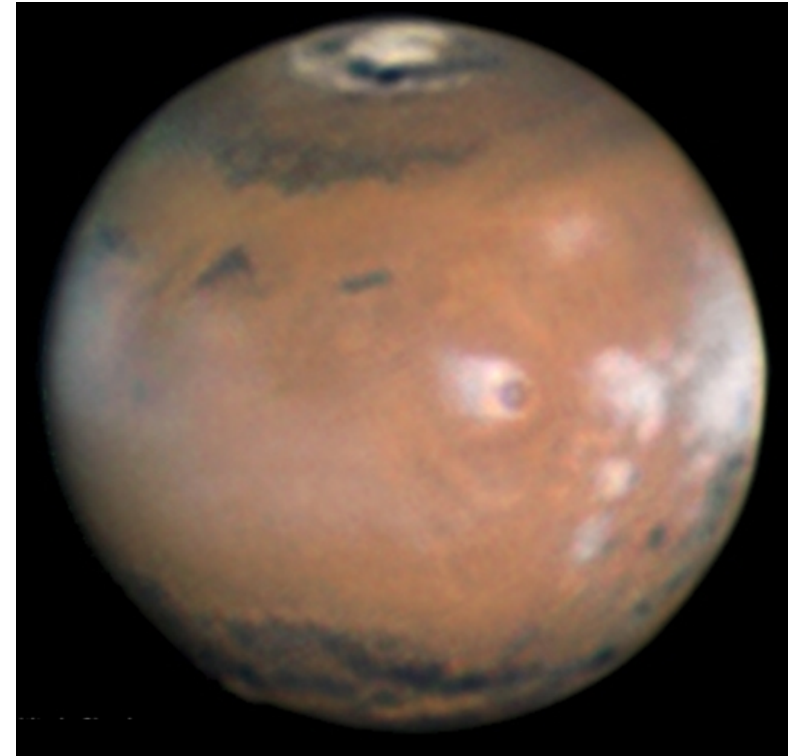
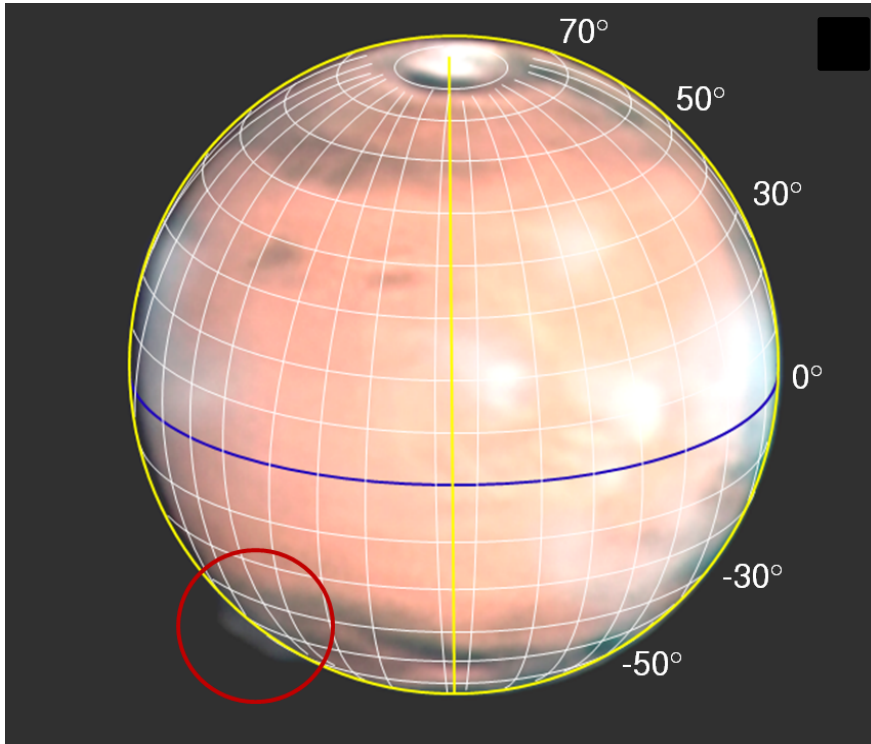
**\* Detached Dust Layers (EDDL)**  
 Transport: updrafts over volcanos, 'rocket dust' convection, waves (tides, gravity), Hadley circulation; Dust storms)



Spiga et al., JGR (2013)

## Plume ('protrusion'): Observations in 2012

Two plume events: (1) 12 – 23 March ; (2) 6 – 16 April



### Observations:

- \* Telescope aperture = 20 - 40 cm
- \* "Lucky imaging" technique
- \* Plate scale  $\sim 0.045$  arcsec/pixel  
( $\sim$ pixel size = 25 km on Mars limb)
- \* Wavelength range  $\sim 450 - 650$  nm,  $> 750$  nm

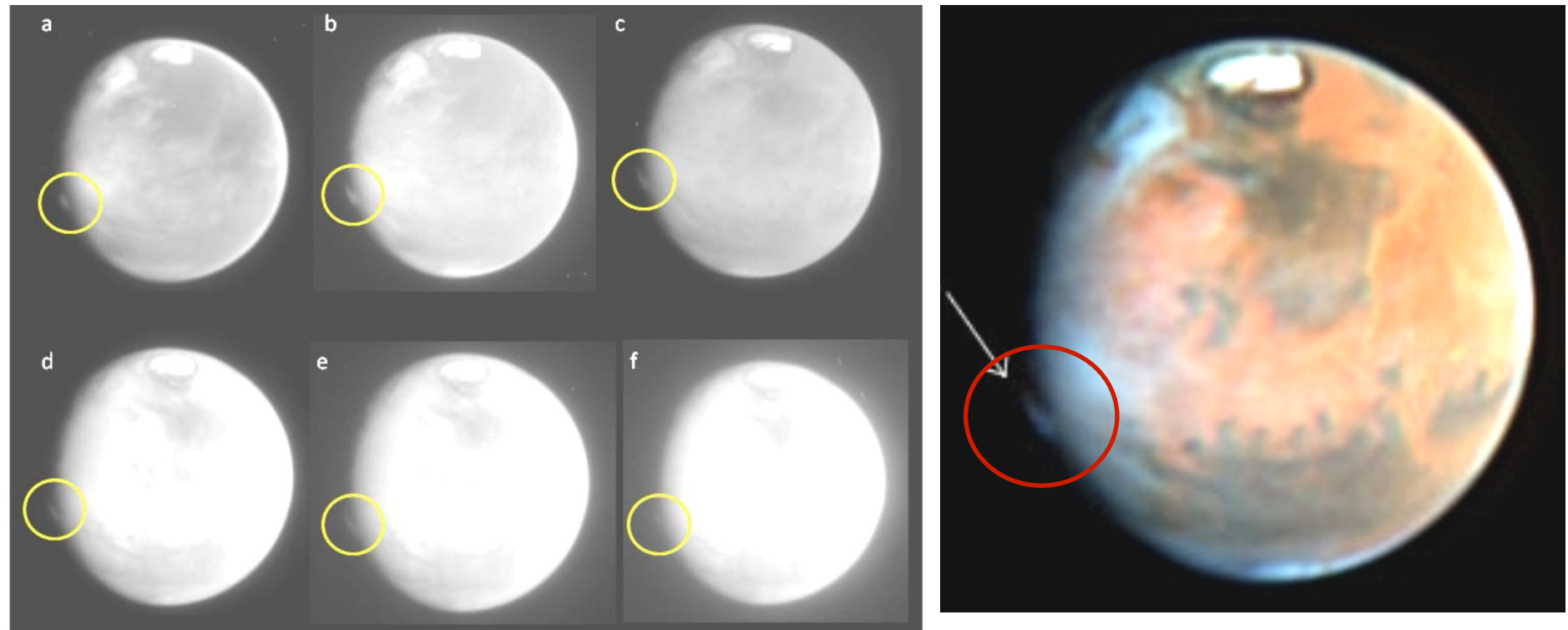
### Observing conditions:

- Mars size =  $13.8''$
- Illumination = 99.6% (phase angle  $\alpha = 7.2^\circ$ )
- $D_{\text{Earth}} = 22^\circ$  ( $D_{\text{Sun}} = 25^\circ$ )  $\rightarrow$  North Pole vision

(a) 18 Independent observers; (b) Experienced observers; (c) 20 Different days; (d) Always same location and Mars local time; (e) Protrusion rotating  $\rightarrow$  **Rule out "an artifact"**



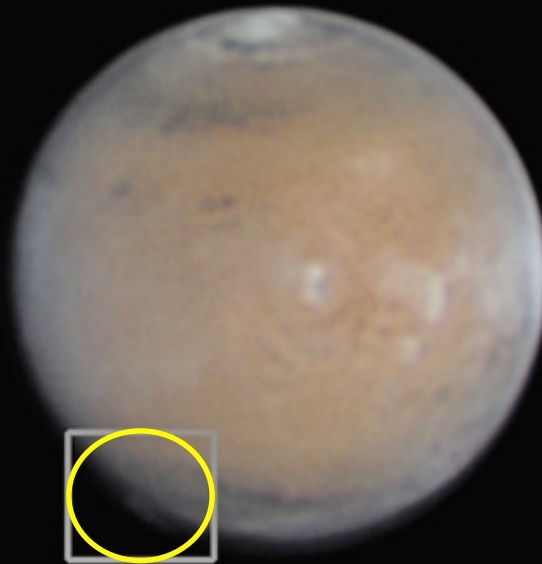
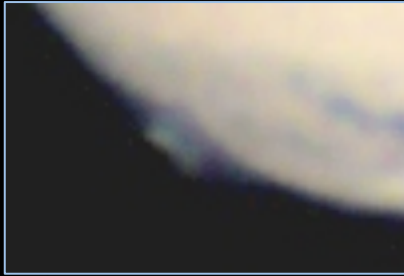
Plume ('protrusion'): 17 May 1997  
Hubble Space Telescope



HST – Wide Field Planetary Camera  
Wavelengths: (a) 257, (b) 409, (c) 502, (d) 589, (e) 673, (f) 1045 nm

# Plume ('protrusion'): Observations in 2012

20 Mar 2012  
(W. Jaeschke)

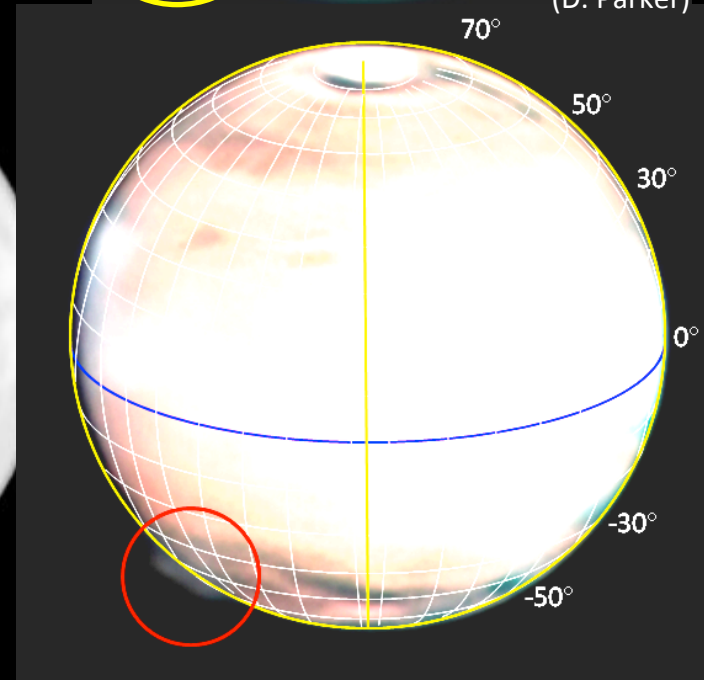
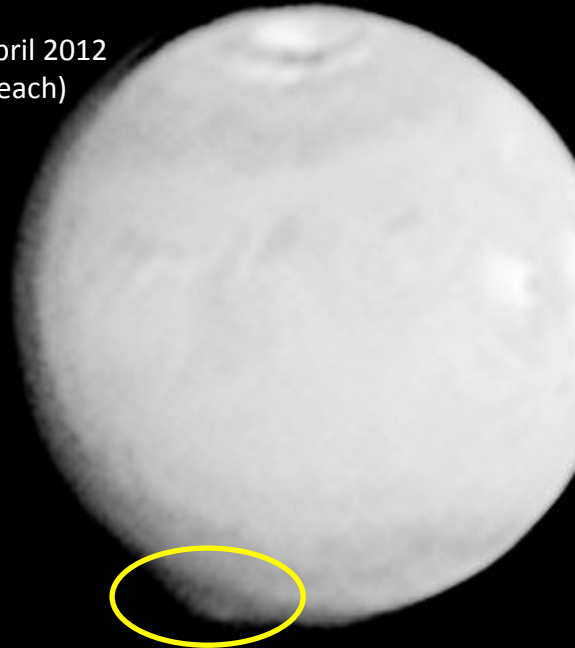


21 Mar 2012  
(D. Parker)

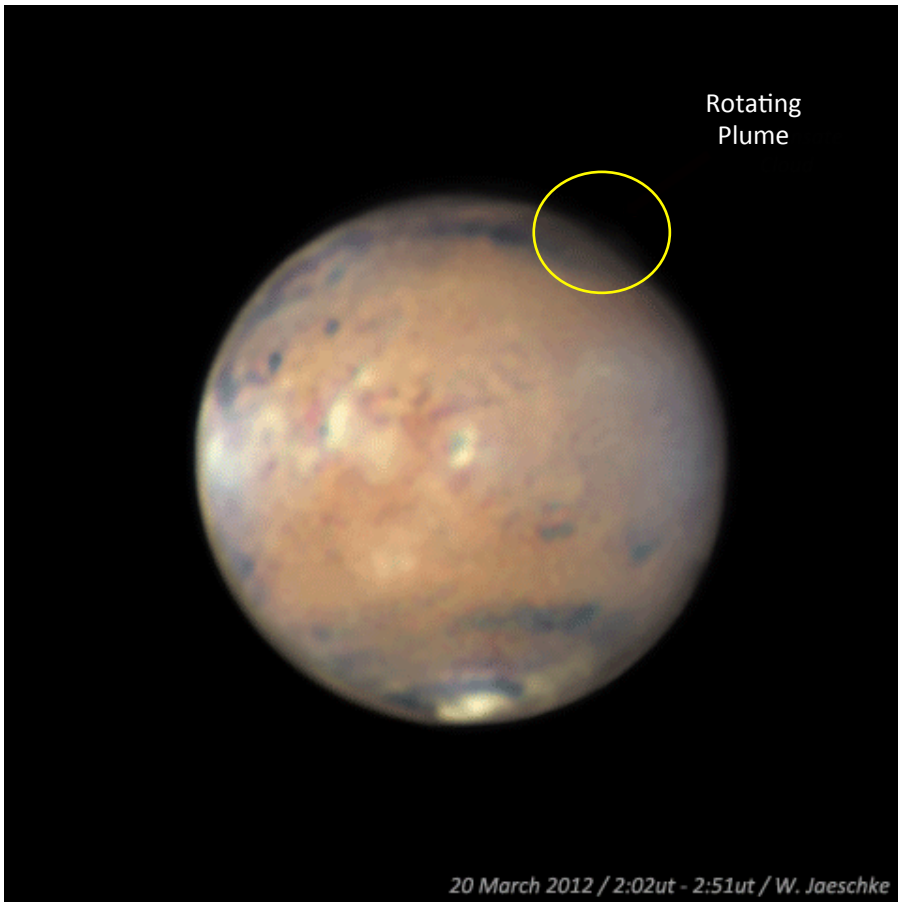
21 Mar 2012  
J. Phillips



13 April 2012  
(D. Peach)



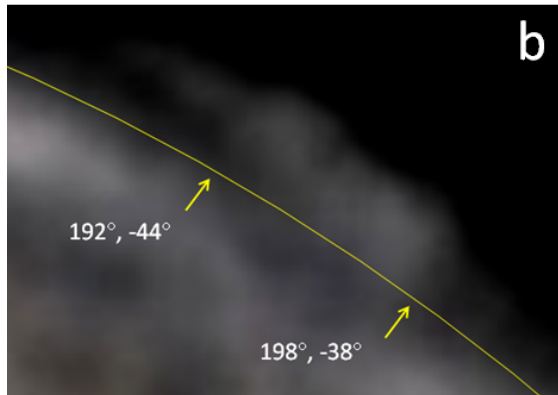
## Plume rotating at morning terminator



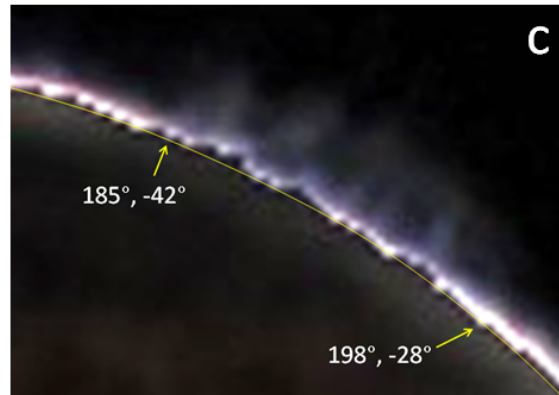
- Overall lifetime of each event  $\geq 10$  days
- Rapid changes with timescales  $< 12$  hrs
- Plume not detected when reaching the evening limb
- Plume not detected when transiting on the CM.

# Plume morphology: Rapid variability

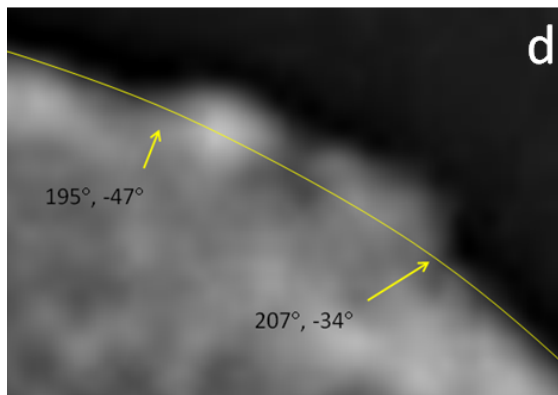
20 March  
(02h 45min)



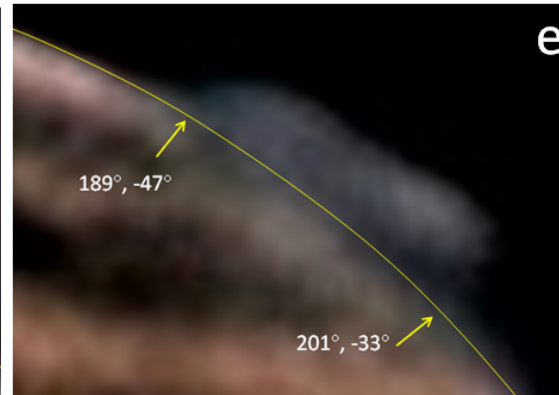
21 March  
(02h 51min)



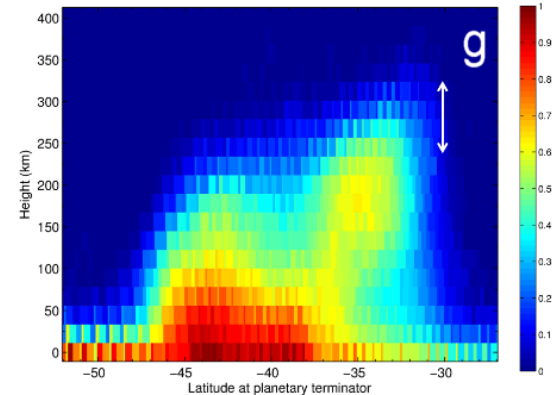
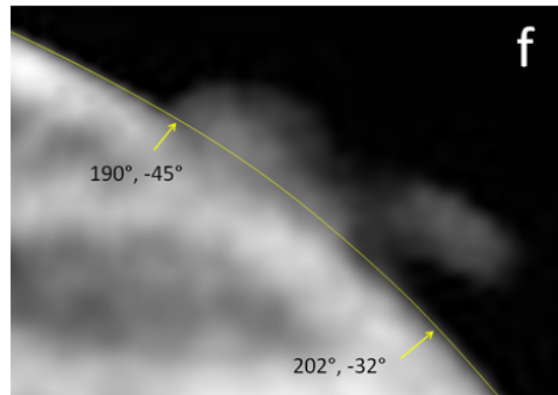
21 March  
(03h 45min)



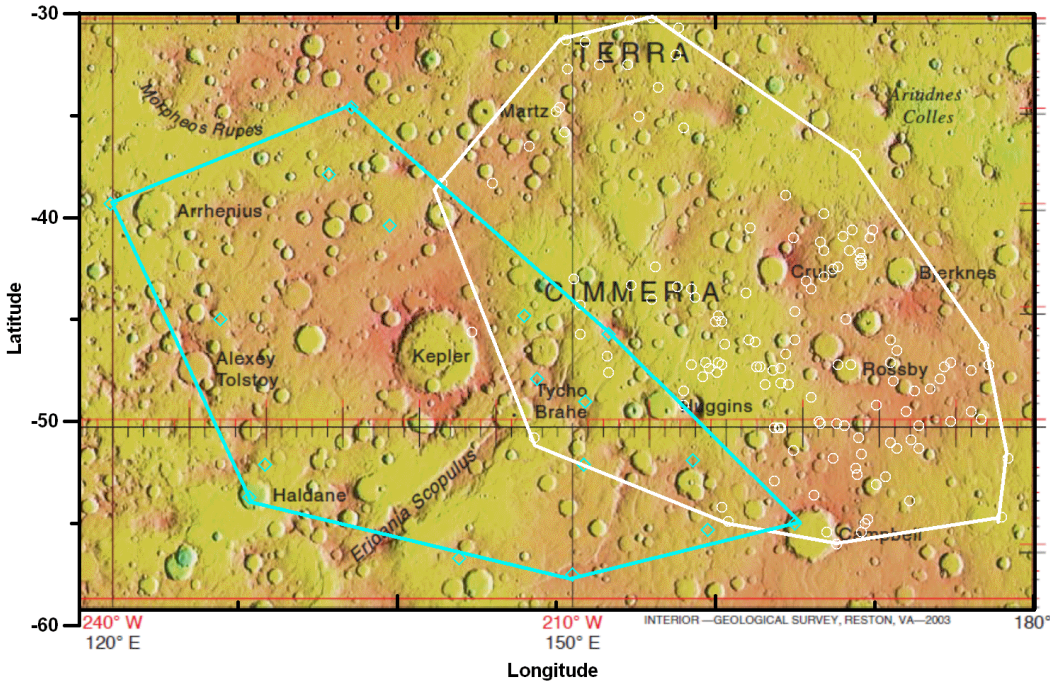
21 March  
(03h 21min)



21 March  
(03h 21min)



## Areographical location of the 1997 & 2012 events



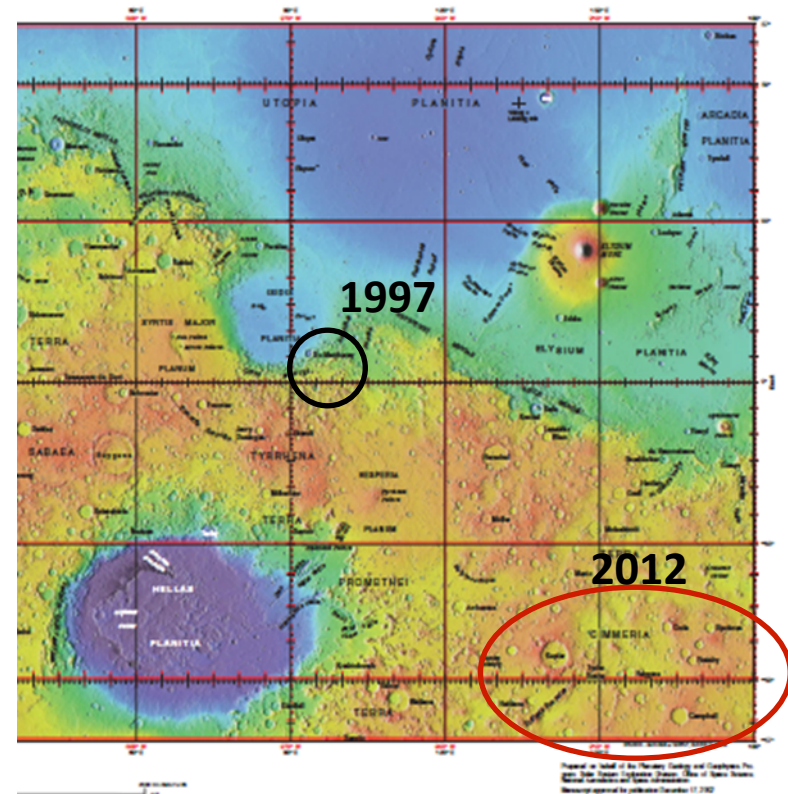
**12 – 23 March 2012**

Latitudes range:  $-35^{\circ} \rightarrow -54^{\circ}$

Longitude (West) range:  $188^{\circ} \rightarrow 225^{\circ}$

Heliocentric longitude:  $L_s = 82.5^{\circ} - 90^{\circ}$

**Early winter (Southern Hemisphere)**



**17 May 1997**

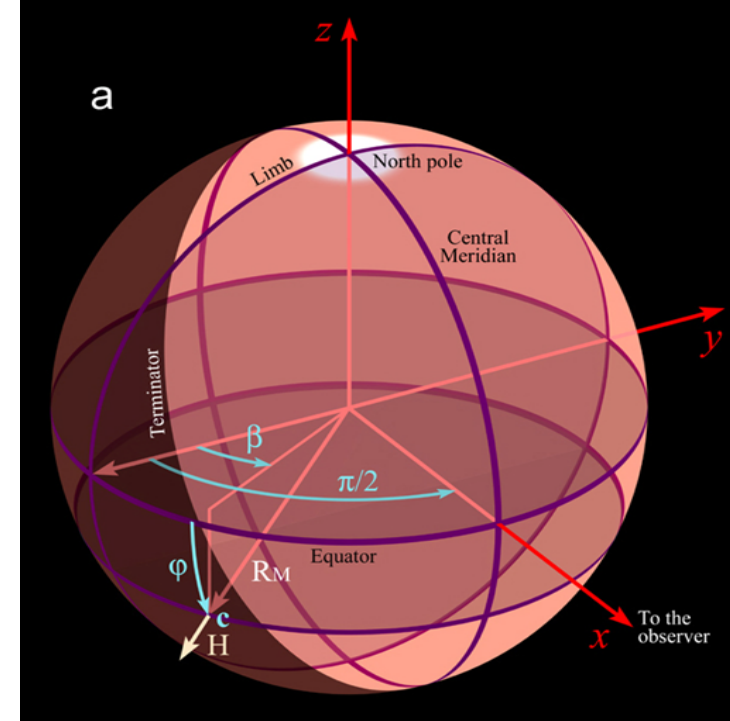
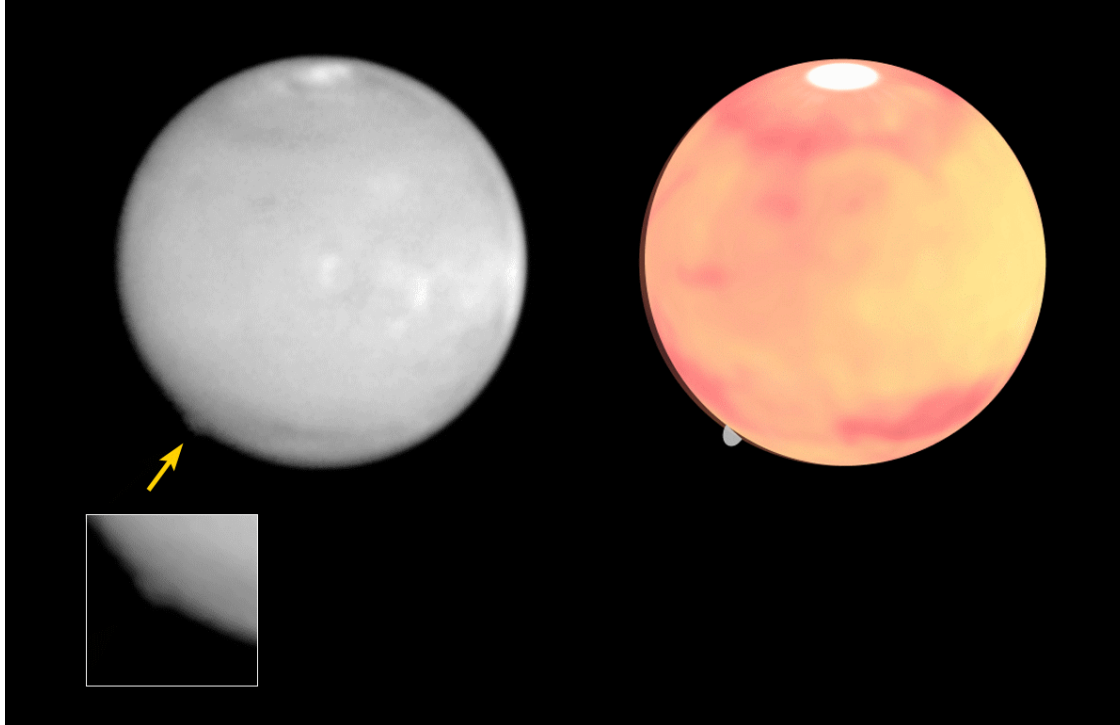
Latitudes range:  $-5^{\circ}S \rightarrow +5^{\circ}N$

Longitude (West) range:  $93^{\circ} \rightarrow 103^{\circ}$

Heliocentric longitude:  $L_s = 119.5^{\circ}$

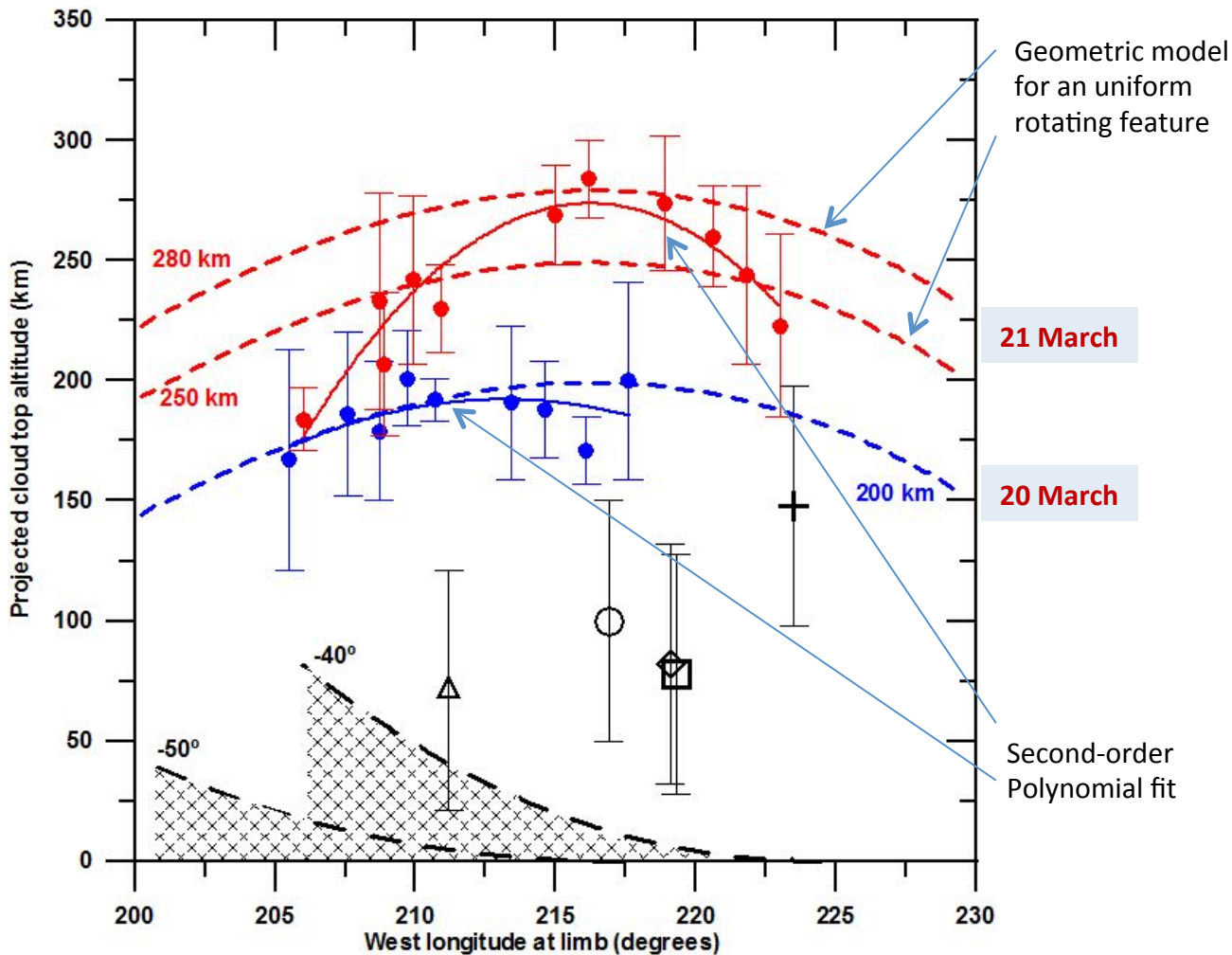
**Summer (Northern Hemisphere)**

# Plume's top altitude: Measurements in rotating frames

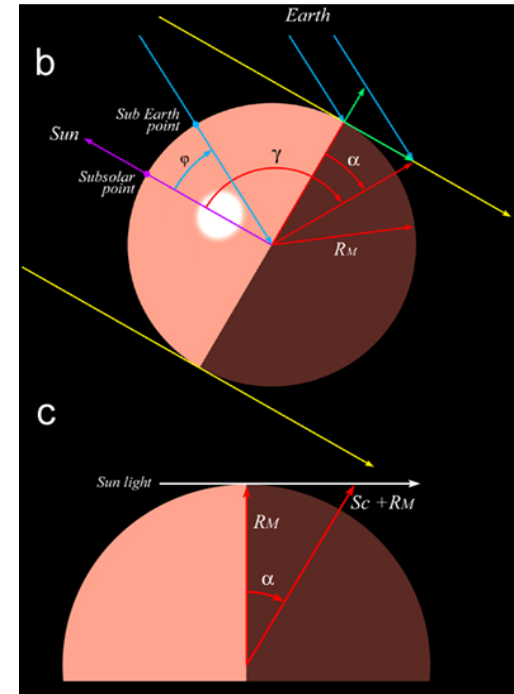


$$\text{Rotating plume model} \rightarrow H = \frac{z'(\beta) + R_M}{\left\{ \cos^2 \varphi_c \cos^2 \beta + (\sin \varphi_c \cos D_E - \cos \varphi_c \sin \beta \sin D_E)^2 \right\}^{1/2}} - R_M$$

# Plume top altitude



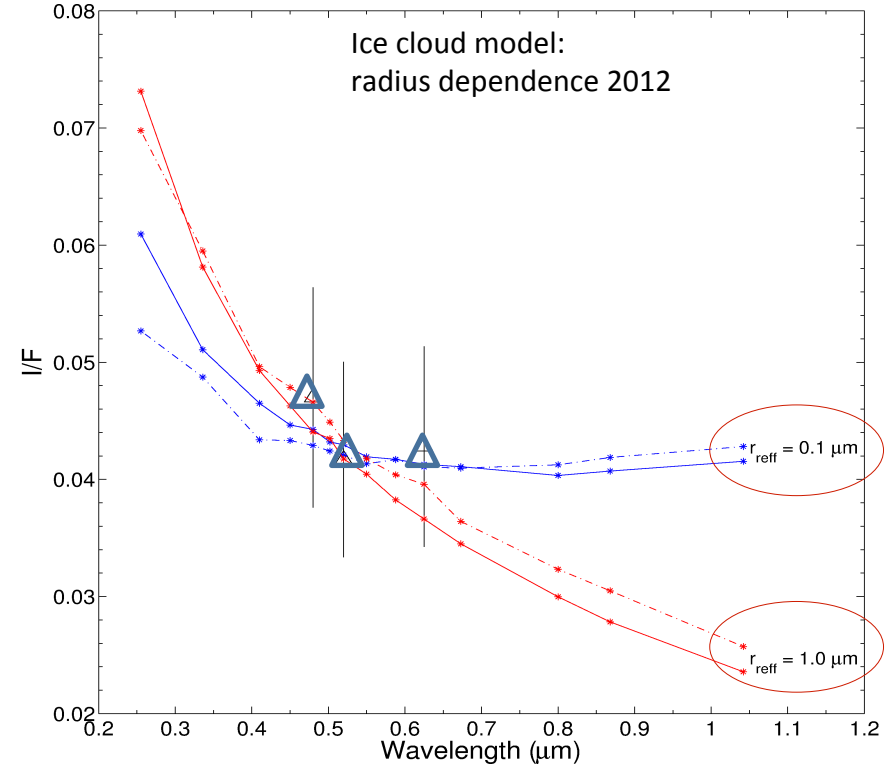
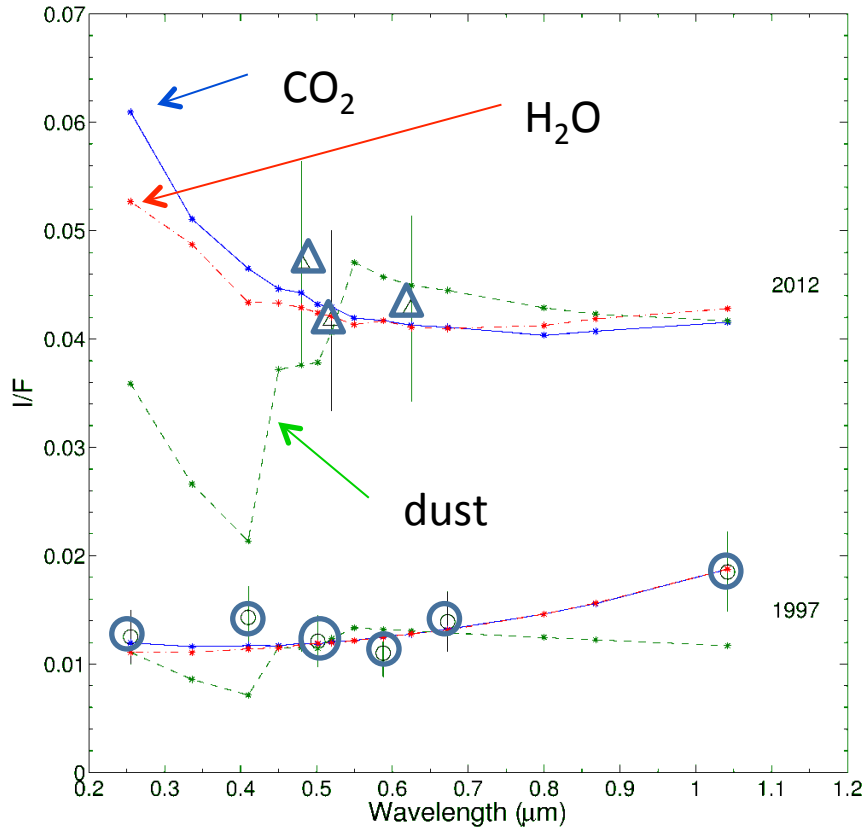
Maximum projected altitude from rotation:  
 20 March  $\rightarrow h = 180 \pm 50$  km  
 21 March  $\rightarrow h = 250 \pm 50$  km



**HST plume altitude:**  
 No rotation available  
 Altitude range:  
 H (min) = 50 km  
 (large horizontal feature illuminated by a grazing Sun)  
 H (max) = 480 km  
 (thin vertical feature under a 90° solar illumination)

# Reflectivity & Cloud model hypothesis

Integrated area photometry in the plume



## Radiative Transfer Model (Mie particles) : Monte Carlo multiple scattering for spherical geometry (\*)

Best fits:  $\text{CO}_2$  or  $\text{H}_2\text{O}$  ice particles

Effective radii =  $0.1_{-0.04}^{+0.1} \mu\text{m}$

Effective variances = 0.1–2.0

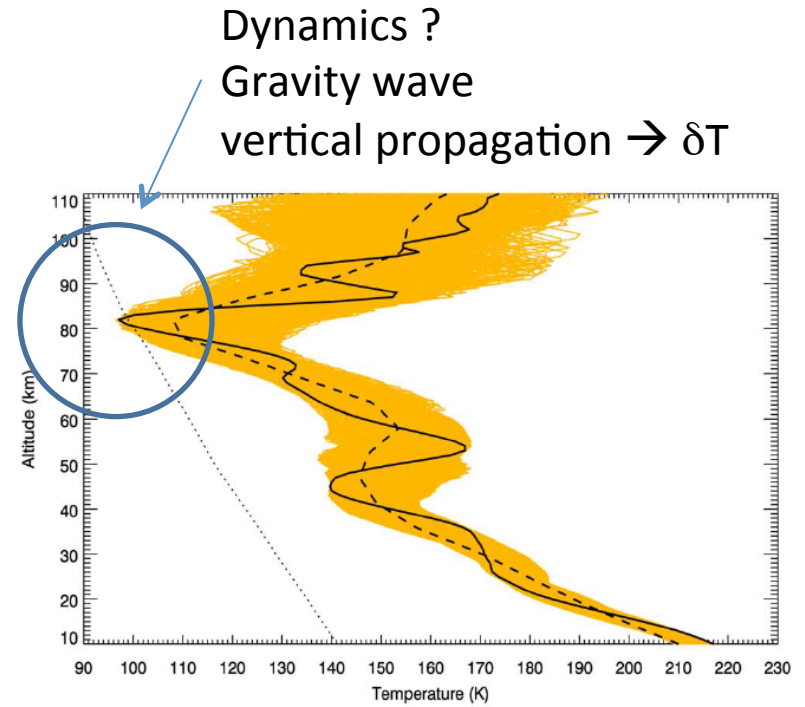
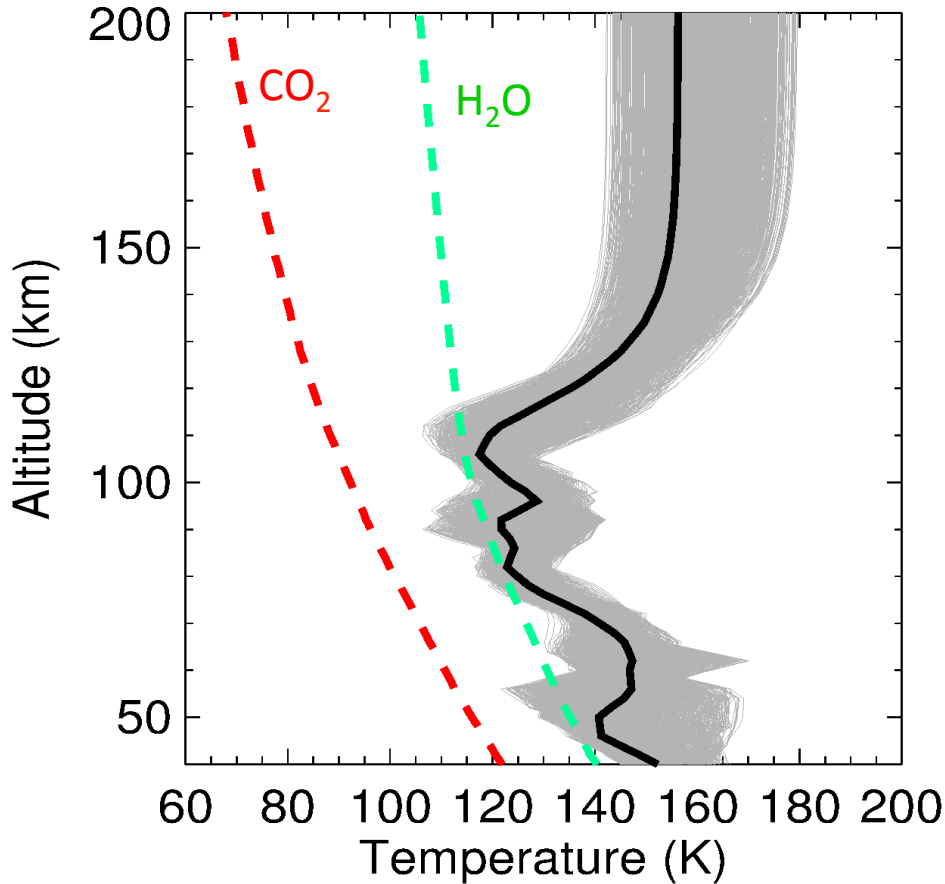
Nadir optical depth  $\tau_N > 0.5$

Number density =  $0.01 \text{ particles cm}^{-3}$

(\*) Garcia Muñoz & Mills, A.A. (2014)



## Ice cloud hypothesis: H<sub>2</sub>O & CO<sub>2</sub>



Spiga et al., JGR (2013)

General Circulation Model →  $T(z)$ ,  $X(z)$  (Gonzalez-Galindo et al., JGR, 2011, 2013)

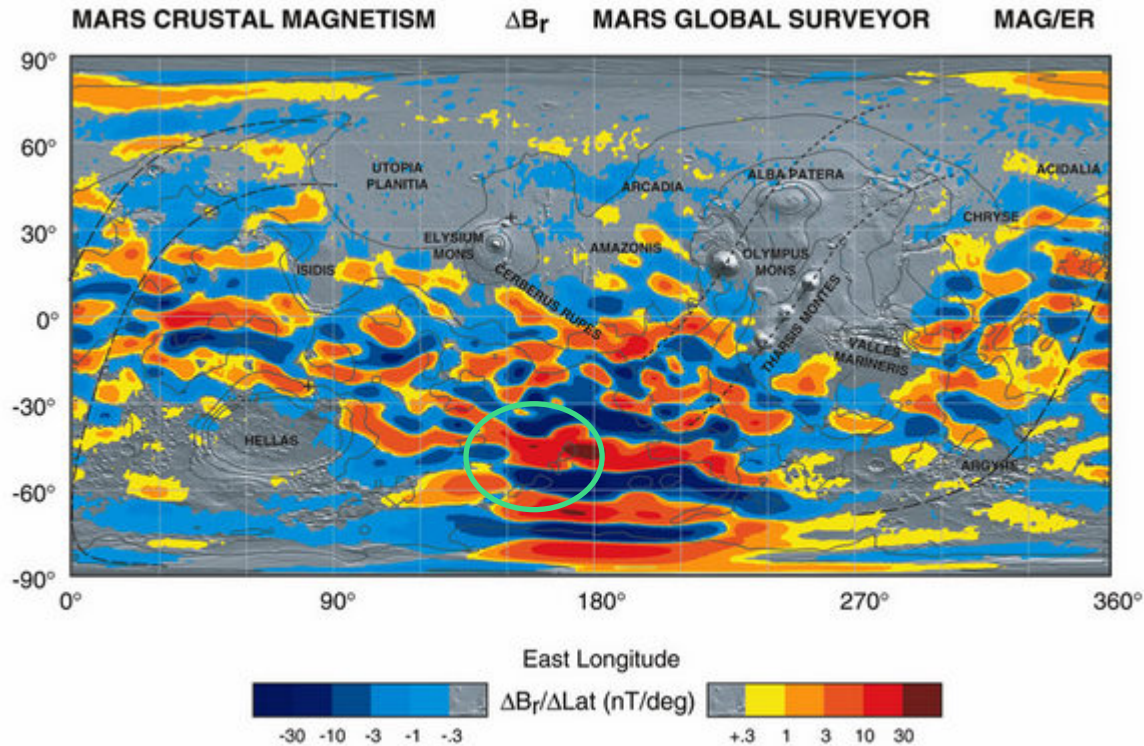
(1) H<sub>2</sub>O –ice condensation

→ anomalously cold thermosphere temperatures (Temperature drop >50 K)

→ unusual increase in the H<sub>2</sub>O mixing ratio from  $10^{-4}$  to complete saturation above 140 km

(2) CO<sub>2</sub> -ice condensation → Temperature drop ~ 100 K above 125 km

# A magnetic related event: Aurora?



Connerney, J. E. P. et al., (2005) Proc. Natl. Acad. Sci. USA, 102, No. 42, 14970-14975.

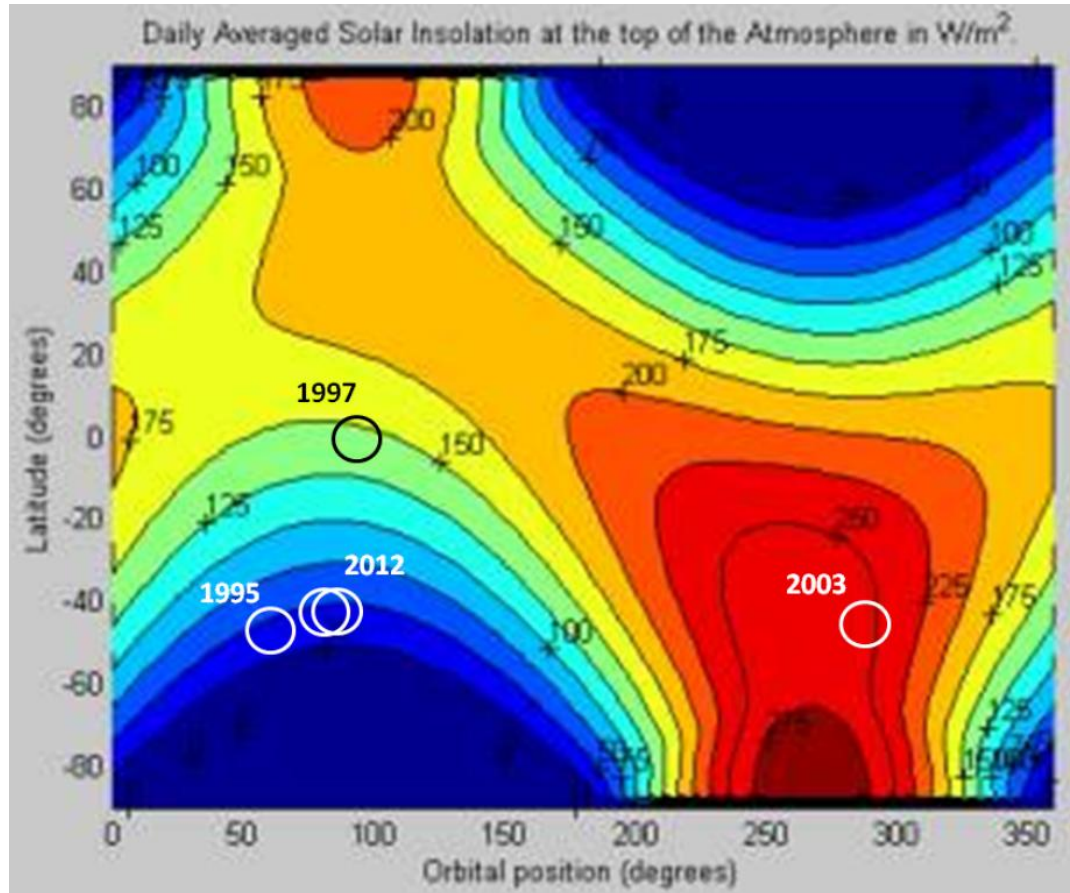
R1599\_1pub

2012 Plume occurred over a large anomaly in the crustal magnetic field (175° East).

Reflectivity  $I/F$  (550 nm) = 0.04  $\rightarrow I_{Aurora}$  (limb)  $\sim 3,600$  MR ( $\gg 1$ -MR nadir emission terrestrial aurorae)  $\gg$  Mars' UV aurora (CO Cameron bands)  $\rightarrow I_{Aurora}$  (limb)  $\sim$  kiloRayleighs.

$\rightarrow$  Exceptional influx of energetic particles over days.

## Seasonal behavior: Events 1997 and 2012 (and in 1995, 2003)



## Conclusions

- **Plumes in 2012:**
- Two events: 10-23 March, 6-16 April (lifetime ~ 10 days)
- Terra Cimmeria (Longitude: 190° - 220°W, Latitude: 45°S)
- Zonal and meridional lengths ~ 500 – 1,000 km → **NOT Detached Layer type**
- Maximum top altitude ~ 200 - 250 km (day to day variability)
- Only visible at morning terminator
- Ls = 82° - 90° (Early winter Southern H.)
  
- **Radiative Transfer Cloud model** → CO<sub>2</sub>, H<sub>2</sub>O ice (no dust)
- $r_{\text{eff}} = 0.1_{-0.04}^{+0.1} \mu\text{m}$ ,  $v_{\text{eff}} = 0.1\text{--}2.0$
- Nadir optical depth:  $\tau_N > 0.5$  (N = 0.01 particles cm<sup>-3</sup>)
- **Requires a large temperature drop or large vapor mixing ratio**
  
- **Aurora** → Crustal magnetic field anomaly.
- Intensity ~ 3,600 MR (~ x1,000 Earth aurora)
- **Requires exceptional influx of energetic particles over days**
  
- **Defy current knowledge of Mars upper atmosphere**
- **Need of continuous limb observations for similar events (particular survey of Terra Cimmeria)**

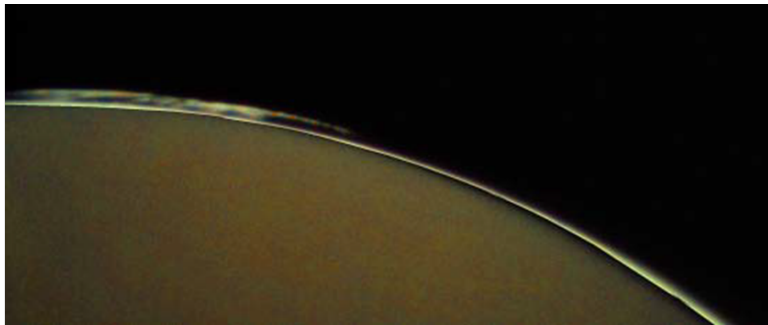
## Future work in progress

### (1) Understand the 2012 events and similar 'not layered' events

Explore models:

- Nature: Clouds? Emission? Both?
- Dynamical mechanisms?
- Seasonal?
- Recurrent? Periodic?
- Areographic dependence?
- Rule out: Internal (geology)? External (influx)?

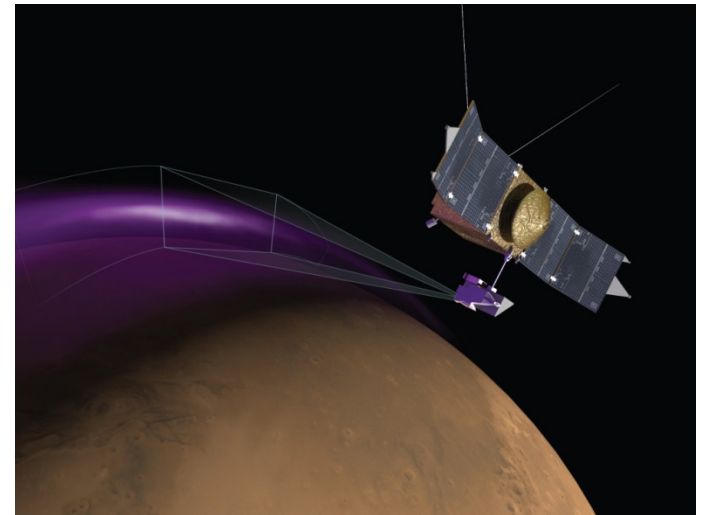
### (2) Continuous limb observations



\* VMC-MEX - 15 Dec.2009

Z(max) ~ 40 km - L ~ 1,000 km

\* HST 1995-1997-1999



MAVEN -IUVS (LPSC, 2015)

Fine dust Z = 150-300 km

UV Aurora (30°-60°N)

GRACIAS POR VUESTRA ATENCIÓN

eman ta zabal zazu



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Aula  
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Gela



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2015

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(\*) Unidad Asociada con el Instituto de Astrofísica de Andalucía-CSIC (2013-14)  
Support by Euskampus-UPV/EHU & funded by Ikerbasque Foundation  
up to 2014 (included). Closed in 2015.



# Research

## **1. Planetary Atmospheres:**

Venus, Mars, Jupiter, Saturn, (Titán), Uranus, Neptune, (exoplanetas)

- General Circulation
- Dynamics and meteorology
- Radiative Transfer (clouds and aerosols)
- Impacts in Jupiter (Saturn)
- (Temperature structure)

## **2. Astronomical and Space instruments:**

- PlanetCam (Vis-SWIR): two-channel lucky imaging astronomical camera
- MEDA for Mars 2020 (NASA)
- MAGIS and JANUS for JUICE (ESA) [launch ~ 2022]
- Software development for image analysis and instrument control

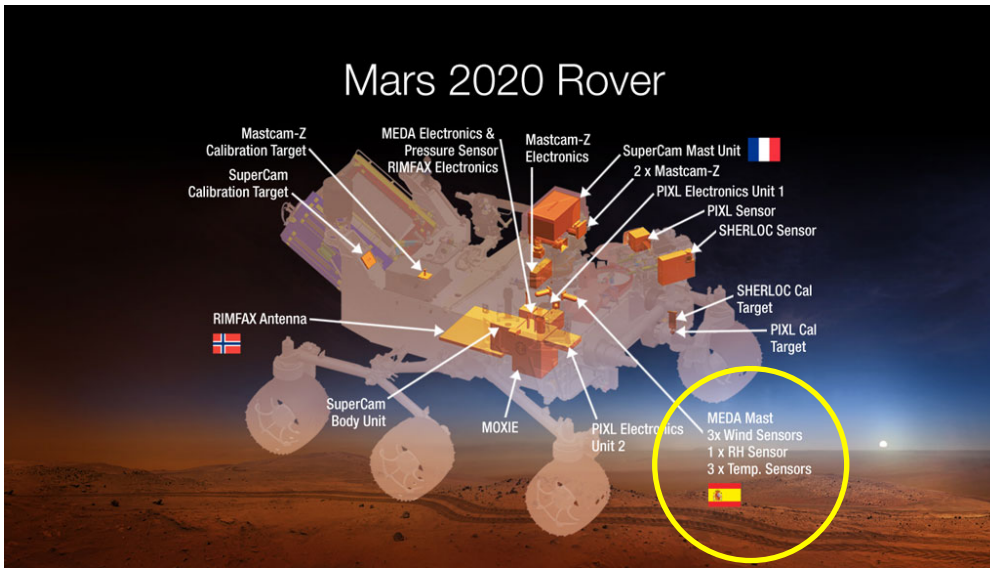
## **3. Data bases management:**

- International Outer Planet Watch (IOPW) & Planetary Virtual Observatory (PVOL)  
→ JUNO mission support (2016)

→ **International collaborations (USA, France, UK, Japan, ...)**

Participation in proposals for ESA – AO

# Instruments Project under development



MEDA: Mars Environmental Dynamics Analyzer

**Moonstruck**  
Space probe will seek signs of life on Jupiter's moons

**Jupiter**  
The largest body in the solar system after the sun

**Europa**  
The size of our moon, surface is almost pure water-ice thought to conceal ocean where life may lurk

**Ganymede**  
Largest moon in solar system, bigger than planet Mercury. Also home to subterranean ocean

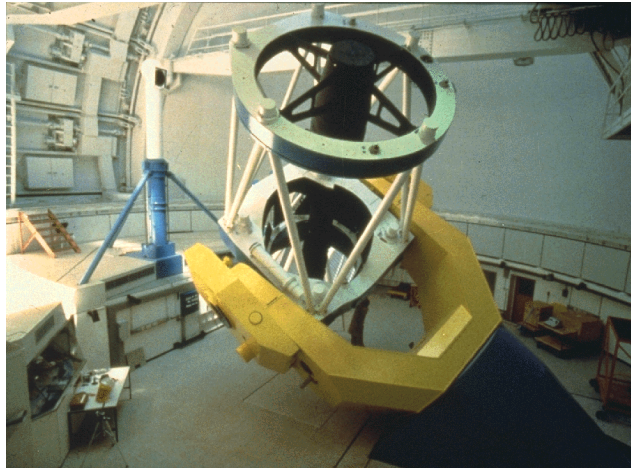
**JUICE**  
Five-ton, billion-euro, solar-powered probe to be launched in 2022, will explore the Jovian moons in unprecedented detail

**Size comparison**

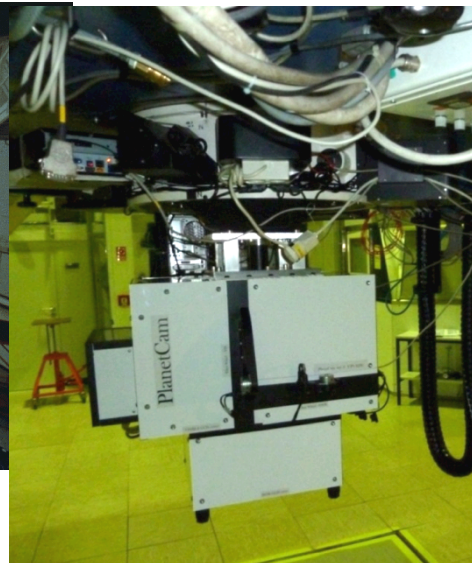
- Jupiter
- Ganymede
- Earth

Diagram labels: Icy shell, Liquid ocean, Deep ice layer

JUICE: Magis & Janus



PlanetCam2 at Calar Alto Obs.



Meade 14" installation at Calar Alto Obs. (lucky imaging)



# Postgraduate formation -Outreach

Castellano | Euskara | English | Contacto



## AULA ESPAZIO GELA

### PATROCINA



Ingeniaritza Goi Eskola Teknikoa  
Escuela Técnica Superior de Ingeniería  
Bilbao



Bizkaiko Foru Aldundia  
Diputación Foral de Bizkaia



Universidad  
del País Vasco Euskal Herriko  
Unibertsitatea

Aula y Master:

<http://www.ehu.es/aula-espazio/>

Master (UPV-EHU):

<http://www.ehu.eus/es/web/cienciaytecnologiaespacial/aurkezpena>

- Master in Space Science and Technology UPV/EHU (2009→, 6 editions)
- Doctorate program (included in Physics UPV/EHU)
- Industry collaboration (Master thesis) & Meetings - Workshops
- Support to students projects
- Outreach activities




 NAZIOARTEKO BIKANTASUN CAMPUSA  
 CAMPUS DE EXCELENCIA INTERNACIONAL

Universidad del País Vasco Euskal Herriko Unibertsitatea

Máster Universitario en Ciencia y Tecnología Espacial

Perfiles | Estudios | Estructura | Investigación | Acceso a la universidad | Áreas temáticas | Servicios | Directorio

Busca en toda la Universidad   
 Búsqueda avanzada

UPV/EHU » Másteres » Másteres oficiales » Máster Universitario en Ciencia y Tecnología Espacial » Información del master » Presentación

**Máster Universitario en Ciencia y Tecnología Espacial**

Información del master

- » Presentación
- » Objetivos y Competencias
- » Programa y profesorado
- » Organización
- » Acceso y matrícula
- » Recursos Materiales
- » Horario y calendario
- » Contacto

Preinscripción y admisión

Verificación, seguimiento y

**Máster Universitario en Ciencia y Tecnología Espacial**

**PRESENTACIÓN**  
 La ciencia y la tecnología espacial representan una de las áreas de vanguardia del conocimiento humano y una de las fuentes actuales más importantes de desarrollo económico, industrial, tecnológico y científico en los países más avanzados. La investigación espacial incluye un amplio abanico de sectores:  
 (1) Investigación e Innovación tecnológica, que incluye el desarrollo de todos los elementos necesarios para la exploración del espacio (plataformas y naves espaciales, telescopios y radiotelescopios de todo tipo en Tierra y



**ENCUENTROS  
ASTROFÍSICA - EMPRESA**  
 BILBAO, 15 de Julio de 2013

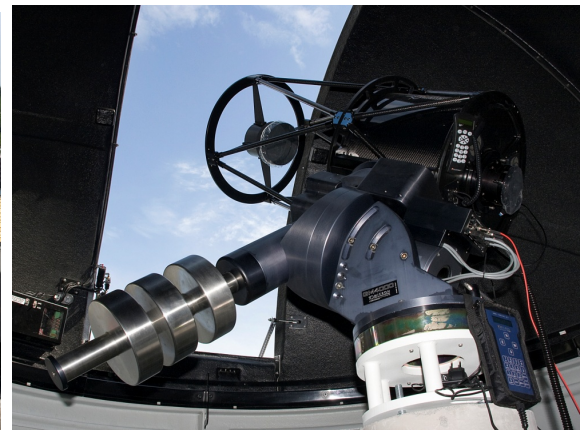
LUGAR: Aula Espazio Gela  
 Escuela Técnica Superior de Ingeniería  
 UPV/EHU, Bilbao



[http://www.ehu.es/aula-espazio/actividades/astrofisica\\_empresa.html](http://www.ehu.es/aula-espazio/actividades/astrofisica_empresa.html)



1. Headquarters:  
Escuela Técnica Superior de Ingeniería  
Computer facilities.
2. Aula Espazio Gela (ETSI)
3. Astronomical Observatory (ETSI)



# Support

1. Project MINECO: AYA36666 (2013-15) + FPI.

**New application 2015**

2. Grupos Consolidados G. Vasco IT765-13 (2013-2015 & 2016-2018)  
Contracts

3. GCP en Unidad Formación e Investigación (UFI) UPV/EHU

~~4. Unidad Asociada con IAA-CSIC (2013-14)~~

~~Ikerbasque support: 33.000 euros/year (\*finished)~~

5. Aula EspaZio Gela.

**Pending new agreement 2015-16 to 2017-18.**

6. Europlanet 2020 Research Infrastructure (H2020)

IP: Louise Thomas (Open University, UK)

Planetary archive

XII REUNIÓN CIENTÍFICA DE LA  
SOCIEDAD ESPAÑOLA  
DE ASTRONOMÍA



**Bilbao**  
18 - 22 Julio de 2016  
Bizkaia Aretoa

INFORMACIÓN:  
<http://www.sea-astronomia.es/SEA2016>

